

TABLE 2.—Vapor pressures at pyrheliometric stations on days when solar radiation intensities were measured.

Washington, D. C.			Madison, Wis.			Lincoln, Nebr.			Santa Fe, N. Mex.		
Dates.	a. m.	p. m.	Dates.	a. m.	p. m.	Dates.	a. m.	p. m.	Dates.	a. m.	p. m.
1918.	mm.	mm.	1918.	mm.	mm.	1918.	mm.	mm.	1918.	mm.	mm.
Aug. 12	19.89	21.28	Aug. 6	14.10	15.11	Aug. 12	13.13	23.52	Aug. 1	7.87	9.83
14	19.23	19.89	13	16.20	16.79	13	17.37	15.11	2	7.87	7.87
15	12.68	12.68	19	11.38	10.59	26	11.38	7.29	3	7.29	6.27
16	13.13	15.11	22	17.96	19.23	29	8.48	14.10	5	6.27	7.29
21	11.38	13.61	24	15.11	12.24	31	7.04	7.57	9	10.59	7.87
23	14.60	16.20	25	13.13	13.13				10	7.87	10.59
			26	13.13	9.83				13	8.48	10.59
									15	7.87	7.87
									17	9.83	7.29
									19	5.79	6.76
									23	9.14	4.75
									26	7.29	9.83
									27	6.02	6.02
									29	7.29	10.59
									30	7.87	8.48

TABLE 3.—Daily totals and departures of solar and sky radiation during August, 1918.

[Gram-calories per square centimeter of horizontal surface.]

Day of month.	Daily totals.			Departures from normal.			Excess or deficiency since first of month.		
	Washing-ton.	Madison.	Lincoln.	Washing-ton.	Madison.	Lincoln.	Washing-ton.	Madison.	Lincoln.
Aug. 1.....	cal. 480	cal. 425	cal. 651	cal. - 3	cal. - 50	cal. 115	cal. - 3	cal. - 50	cal. 115
2.....	467	545	627	- 15	72	94	18	22	209
3.....	477	573	627	- 4	102	96	22	124	305
4.....	314	286	603	-166	-183	135	-188	- 59	440
5.....	542	495	614	63	28	38	-125	- 31	528
6.....	533	588	571	56	124	48	- 69	93	576
7.....	476	480	215	2	27	-304	67	120	272
8.....	536	375	294	65	- 84	-222	- 2	36	50
9.....	463	538	306	- 5	81	-206	7	117	-156
10.....	296	475	574	-169	21	65	-176	138	- 91
11.....	388	258	415	- 79	-193	- 90	-255	- 55	-181
12.....	503	552	613	44	104	111	-211	49	- 70
13.....	440	509	431	- 16	64	- 67	-227	113	-137
14.....	503	573	287	50	130	-208	-177	243	-345
15.....	609	411	447	159	- 29	- 44	- 18	214	-389
16.....	575	77	521	128	-361	33	-110	-147	-356
17.....	126	121	315	-318	-314	-169	-208	-401	-525
18.....	84	499	520	-356	67	39	-564	-394	-486
19.....	525	593	525	88	163	57	-476	-231	-429
20.....	543	429	360	109	2	-115	-367	-229	-544
Decade departure.....							- 191	-367	-453
21.....	593	365	590	162	- 60	118	- 205	-280	-426
22.....	437	532	405	9	110	- 64	-196	-179	-490
23.....	500	243	389	75	-177	- 77	-121	-356	-567
24.....	574	558	564	152	141	101	31	-215	-406
25.....	500	504	474	80	90	14	-111	-125	-452
26.....	400	565	599	- 18	154	131	93	29	-321
27.....	412	478	521	- 4	70	68	80	99	-255
28.....	89	283	358	-325	-122	- 95	-236	- 23	-350
29.....	372	527	542	- 40	125	92	-276	102	-258
30.....	422	297	142	12	-102	-306	-264	0	-564
31.....	338	543	614	- 70	152	168	-334	152	-396
Decade departure.....							+ 33	+381	+148
Excess or deficiency since first of year.....	{Or.-cal.						-1,266	+938	+276
	{Per cent.						- 1.3	+1.0	+0.2

The smoothed percentages thus obtained have been plotted as dots on figure 1.

Reference to Table 1 shows that for the period December, 1882–August, 1892, data were available for a single station only, and from September, 1892, to December, 1895, from two stations only. In consequence the monthly values are much more scattering previous to 1896, and especially previous to 1893, than subsequent to these years.

Furthermore, all the radiation measurements previous to 1905 were made at noon only. It follows from the Bouguer equation³ that the percentage of depletion due to diminished transmissibility of the atmosphere, or the percentage increase due to increased atmospheric transmission, would be greater during the winter months, when the sun's zenithal distance is at a maximum, than in the summer months, when it is at a minimum. There is evidence of this annual variation in the radiation intensities, particularly in the early part of the curve.

obtained from stations in Europe. In consequence, the data are not so representative of conditions in the Northern Hemisphere as a whole, but are more influenced by local atmospheric conditions. Aside from the depressions noted, the smoothed monthly radiation intensities show remarkably uniform values. It is because the values obtained during the great depressions have been included in the means that the intensities measured under normal conditions generally fall between 102 and 104 per cent.

The following are the sources of the solar radiation data that have been utilized:

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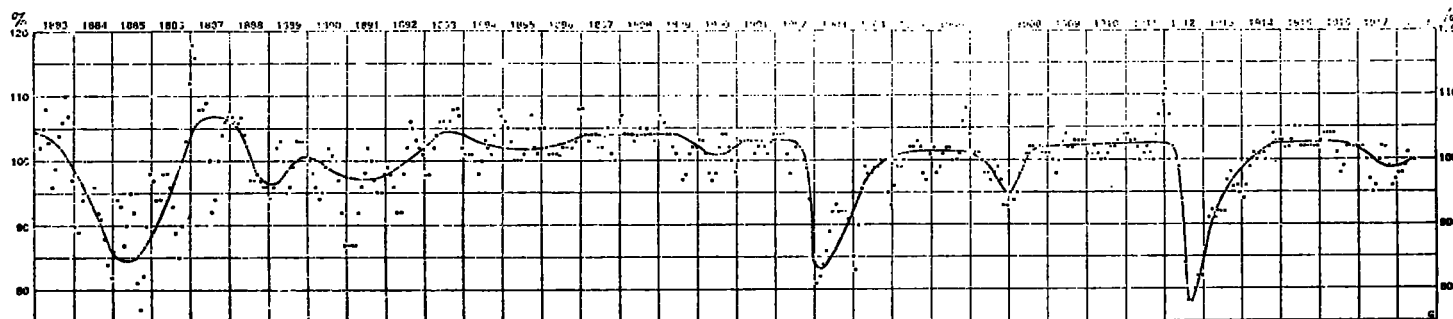


FIG. 1.—Monthly averages of solar radiation intensity at earth's surface, expressed as percentages of the monthly normals.

This has been taken into account in drawing a free-hand curve through the plotted values.

As previously shown,⁴ the three great depressions in the curve of figure 1 follow the great eruptions of Krakatoa in the summer of 1883; of Pelée, Santa Maria, and Colima, in 1902, and of Katmai in June, 1912. In each case there has been complete recovery—in the case of Krakatoa, after 34 months, of Pelée after 17 months, and of Katmai, after 21 months. There are minor depressions in 1888–89 and in 1907–8 that have not been accounted for definitely, and also one in 1916–17 that may be due to the fact that since the outbreak of the European war it has been necessary to substitute data from stations in the United States for that previously

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MADISON, WIS. Bulletin of the Mount Weather Observatory, 5:173–183. Monthly Weather Review, 1916, 44:3, 9–12. [Monthly thereafter.]

SANTA FE, N. MEX. Monthly Weather Review, 1915, 43:439–443, 591. [Monthly thereafter.]

LINCOLN, NEBR. Monthly Weather Review, 1916, 44:4–8. [Monthly thereafter.]

³ $I = I_0 a^m$, where I_0 is the radiation intensity outside the atmosphere, a is the transmission coefficient of the atmosphere, or the proportional part of the radiation transmitted when the sun is in the zenith, m is the length of the path of the sun's rays through the atmosphere in terms of the length of the path for zenithal sun, and I is the solar radiation intensity measured at the surface of the earth. This equation is strictly true for homogeneous rays only.

⁴ Kimball, H. H. loc. cit. Humphreys W. J. loc. cit.